

Resolution Dependence of Storm Time Response using WACCM-X/GAMERA

Hanli Liu¹, Kevin Pham¹, and Francis Vitt^{1,2}

1. NCAR/HAO
2. NCAR/ACOM

WACCM-X and GAMERA

- WACCM-X:

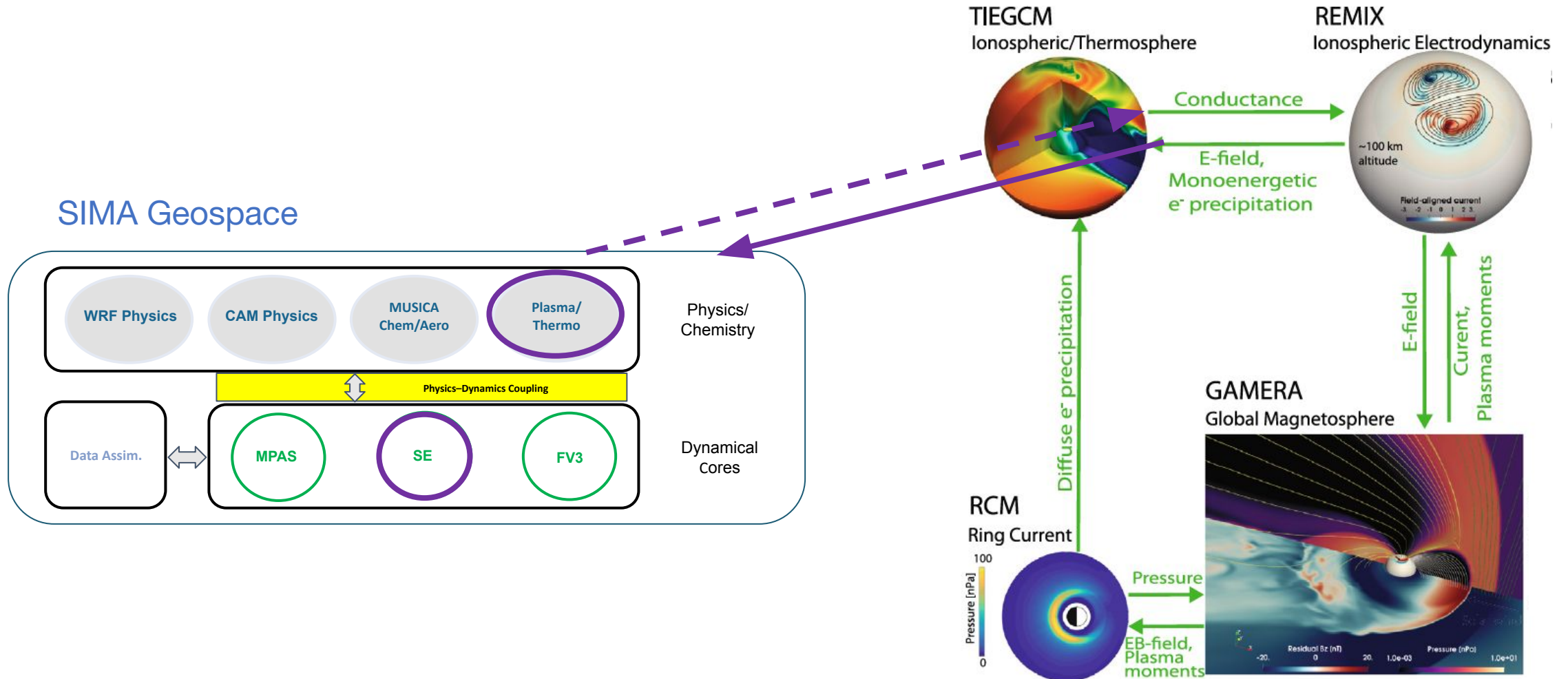
- WACCM-X species dependent spectral element dynamical core on a cubed-sphere grid
- One of the atmosphere components of CESM
- Geospace component of SIMA for space weather applications
- Regular resolution configuration (NE30/L130): ~ 100 km/0.25 scale height
- High resolution configuration (NE120/L273): ~ 25 km/0.1 scale height
- High latitude electric potential specified by GAMERA

- GAMERA:

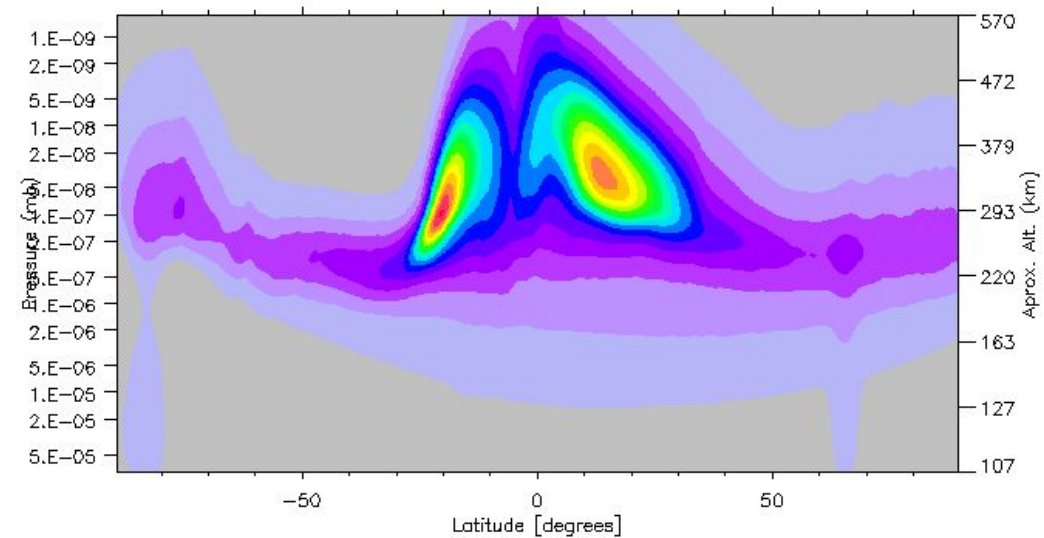
- Grid-Agnostic MHD for Extended Applications is a new magnetosphere MHD code.
- Part of the Multiscale Atmosphere Geospace Environment (MAGE) model, developed by the NASA DRIVE Center: Center for Geospace Storms (CGS).
- Oct-resolution (~ 0.5 deg) simulation for August 24 2005 geomagnetic storm.

Space Weather Application: Storm Time Response

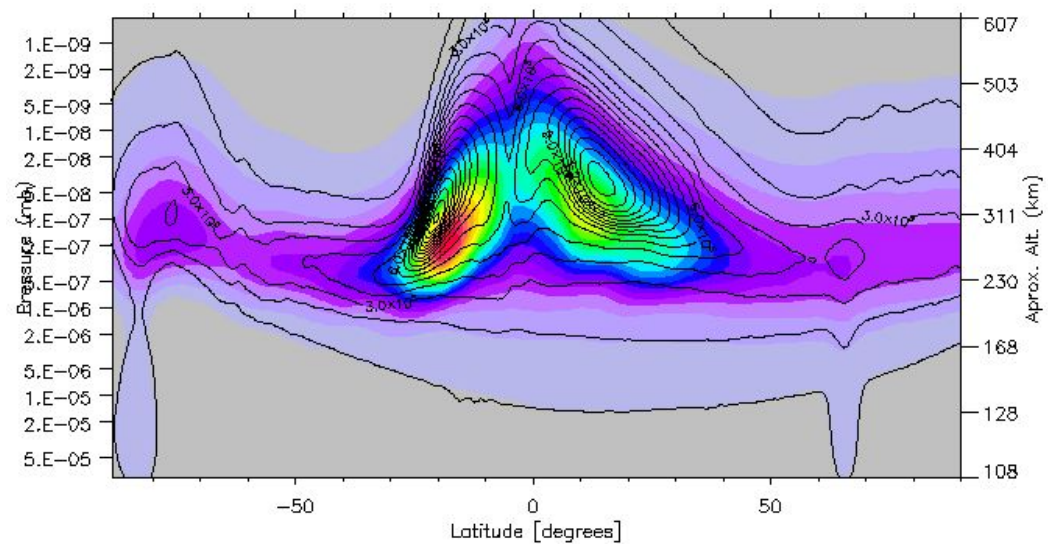
Multiscale Atmosphere Geospace Environment (MAGE) v1.0



EDens [cm^{-3}], 24Aug 1 00:00, lon 240.50000

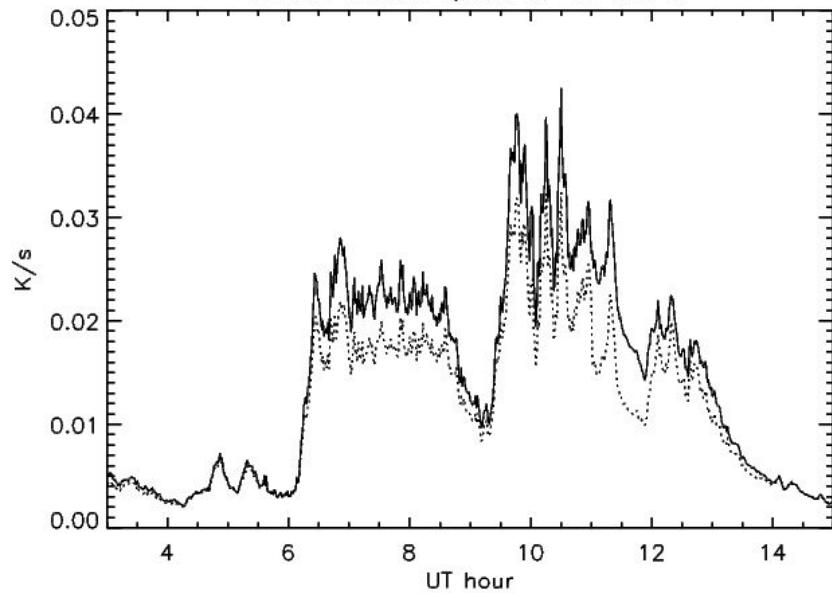


EDens [cm^{-3}], 24Aug 1 00:00, lon 240.46875

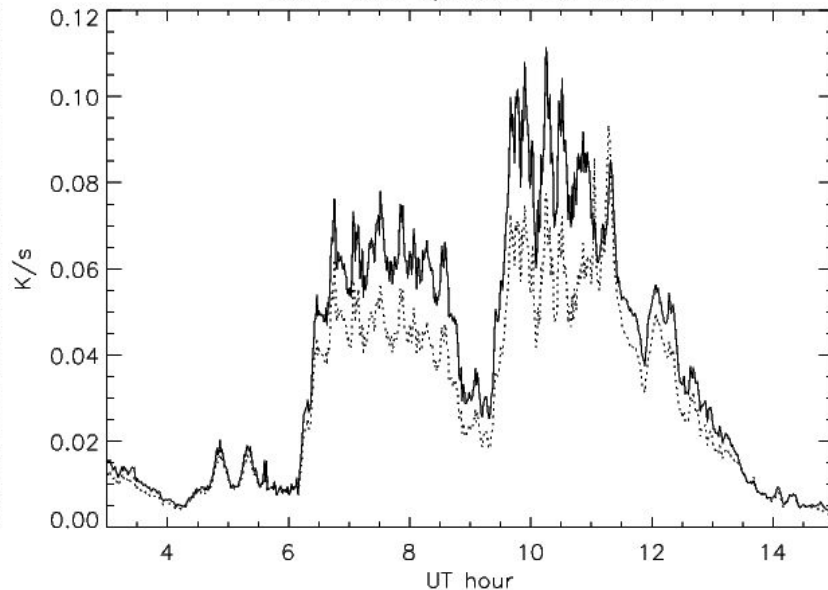


Total Joule Heating at Peak Altitude

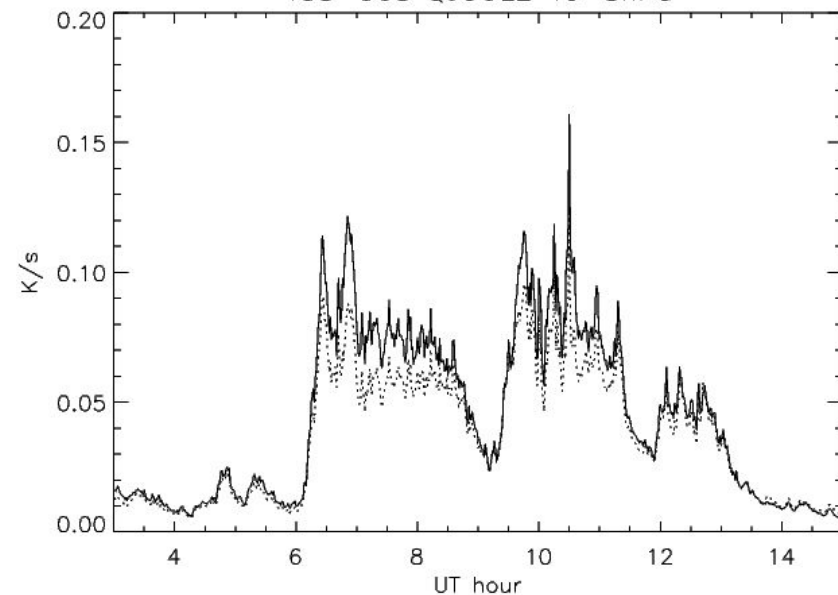
Global mean QJOULE $4e-8hPa$



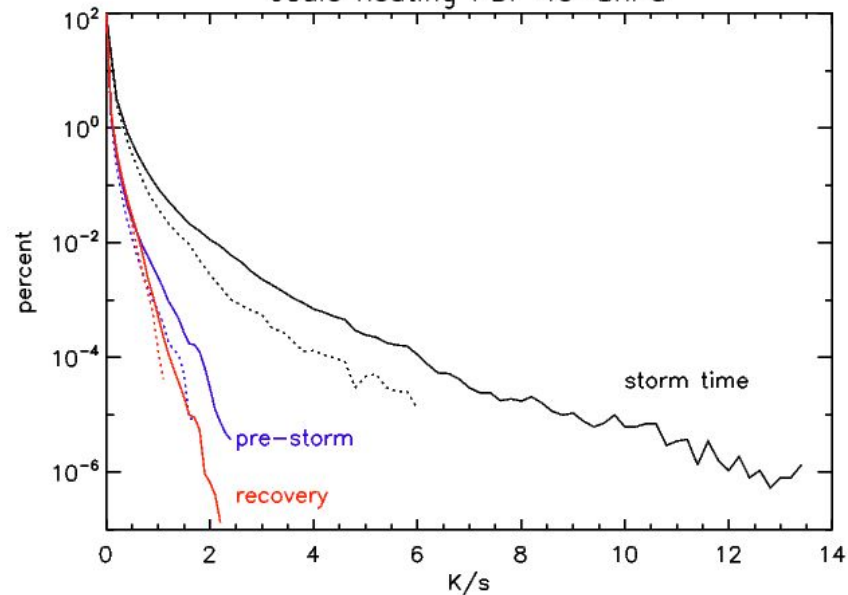
45N-90N QJOULE $4e-8hPa$



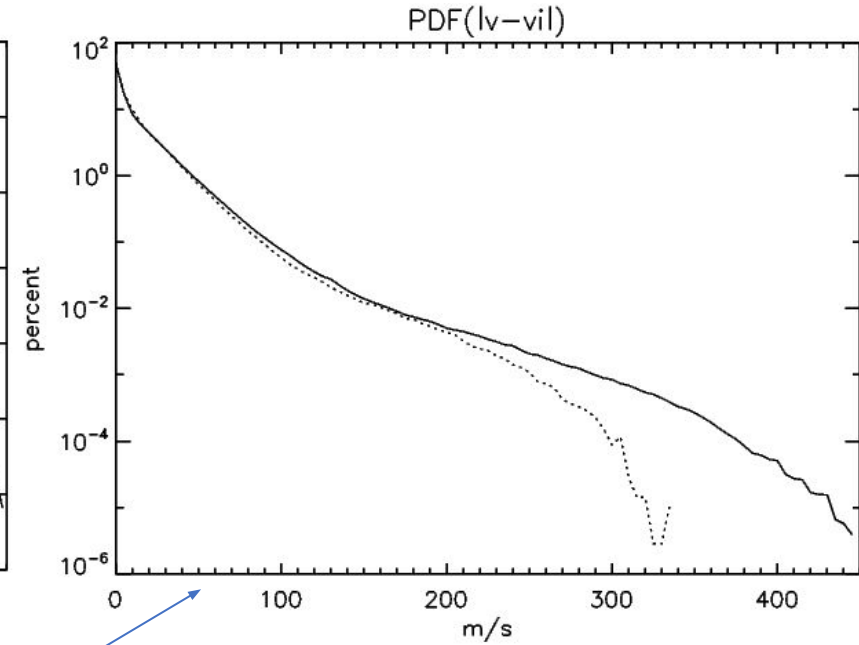
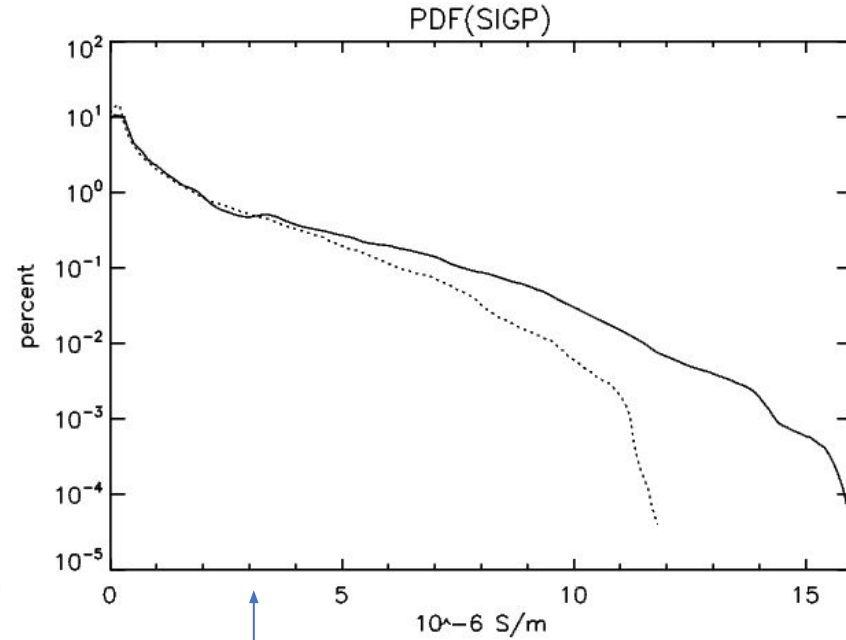
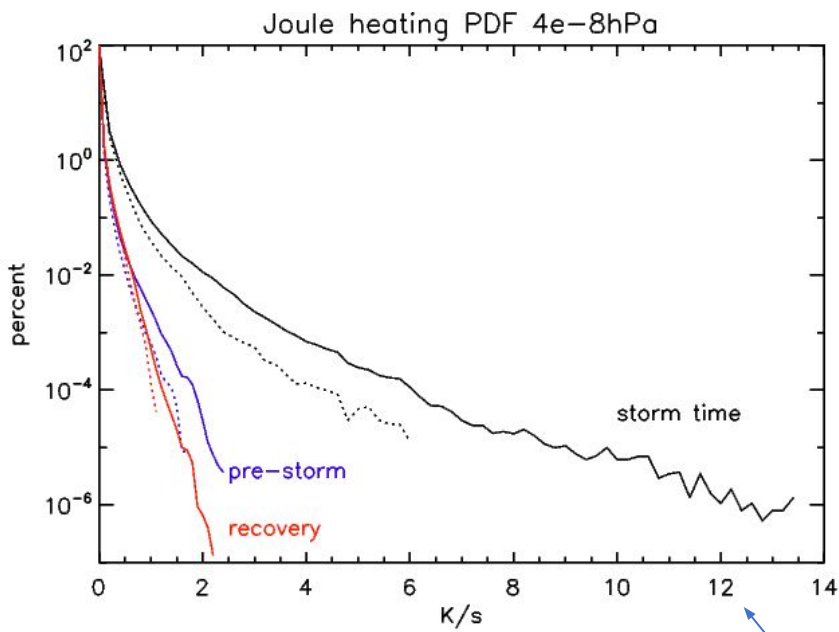
45S-90S QJOULE $4e-8hPa$



Joule heating PDF $4e-8hPa$



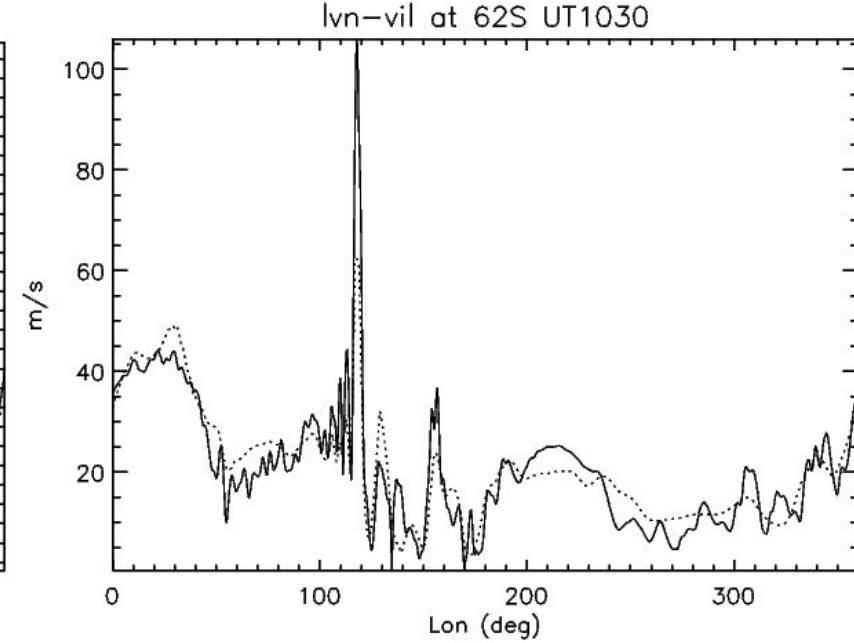
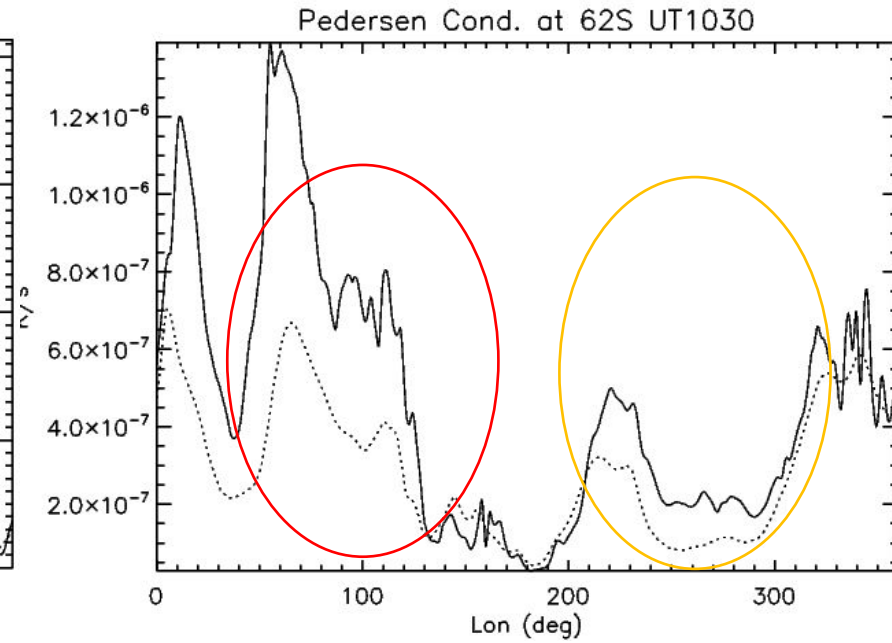
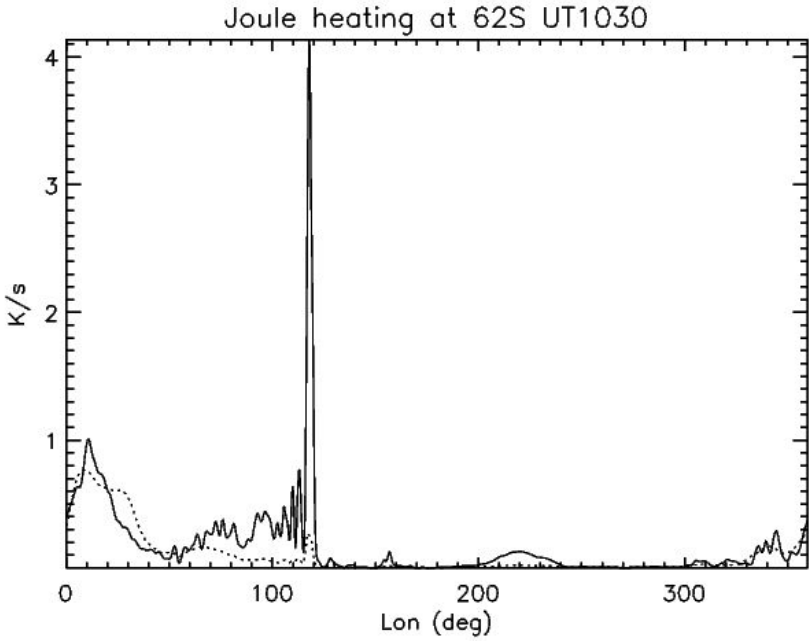
Statistics of Joule Heating



$$Q_J = \frac{\sigma_p B^2}{\rho} (\mathbf{V}_{E \times B} - \mathbf{V}_{n \perp})^2$$

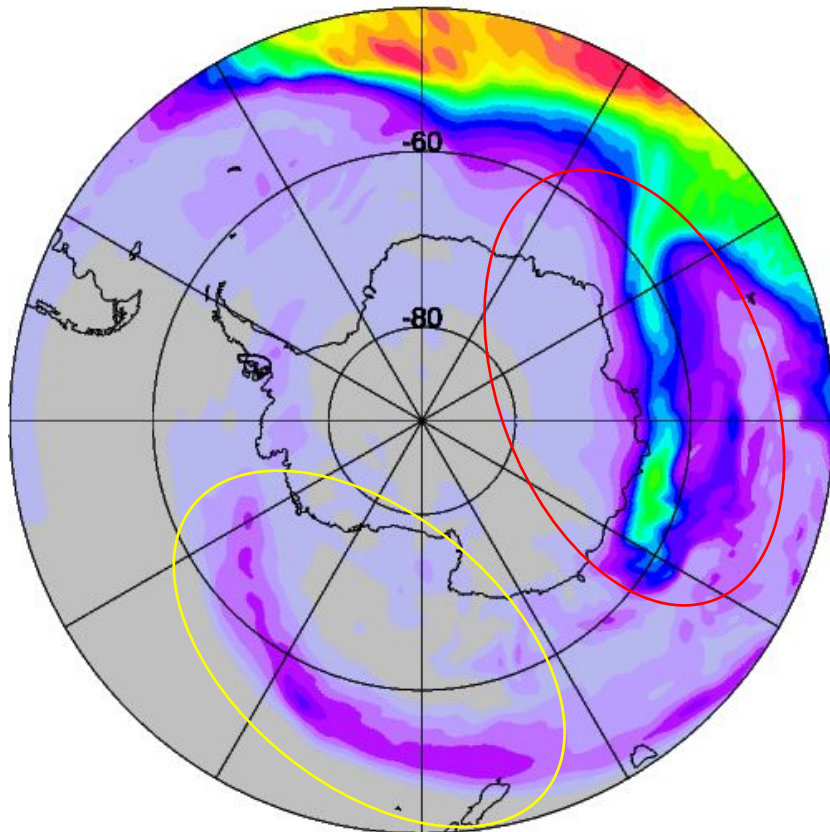
Blue arrows point from the equation to the corresponding x-axes of the three plots above: one to the 'K/s' axis, one to the '10⁻⁶ S/m' axis, and one to the 'm/s' axis.

Joule heating: Conductivity and Wind/Drift Diff.

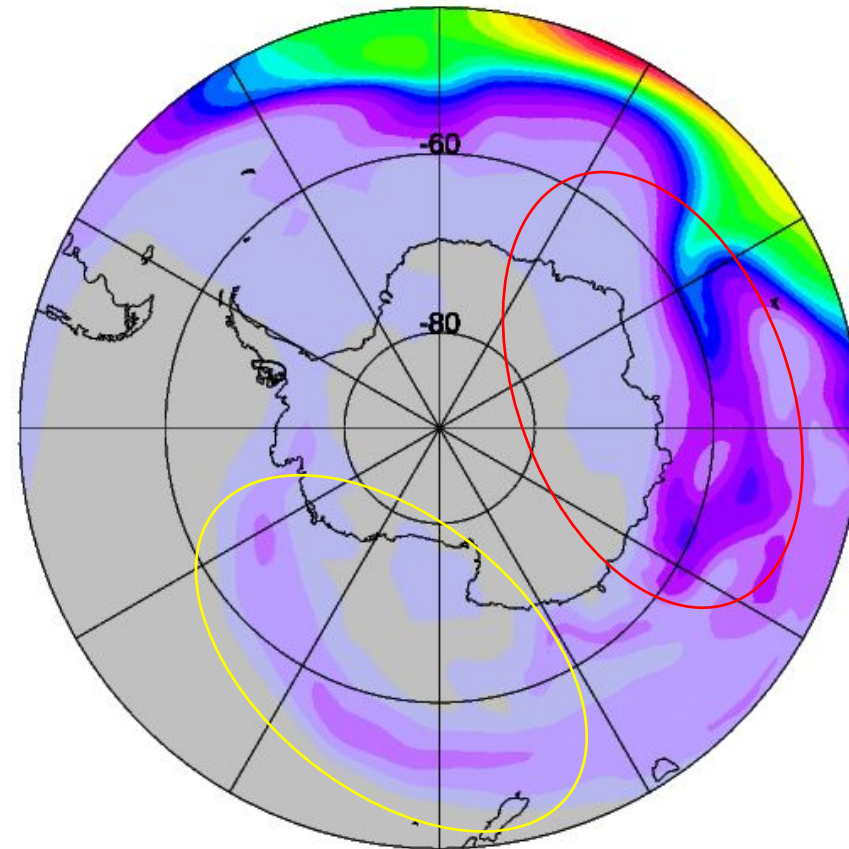


Electron Density:

EDens [cm⁻³], ca. 3.9391303e-08 hPa, 24Aug2005 10:30

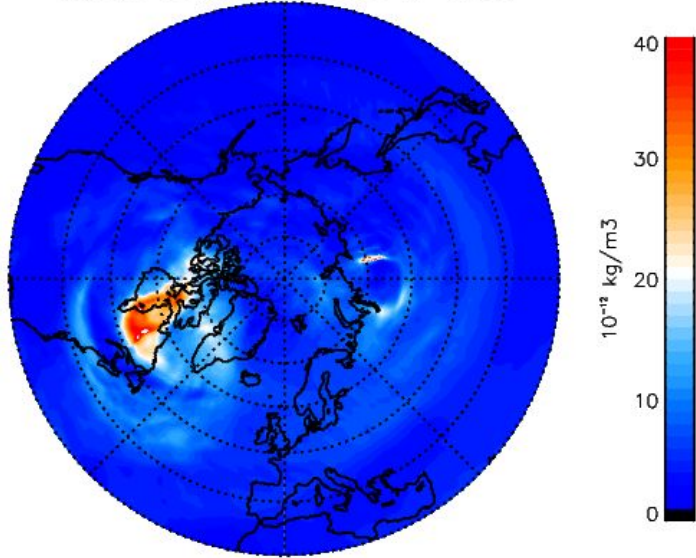


EDens [cm⁻³], ca. 3.6503215e-08 hPa, 24Aug2005 10:30

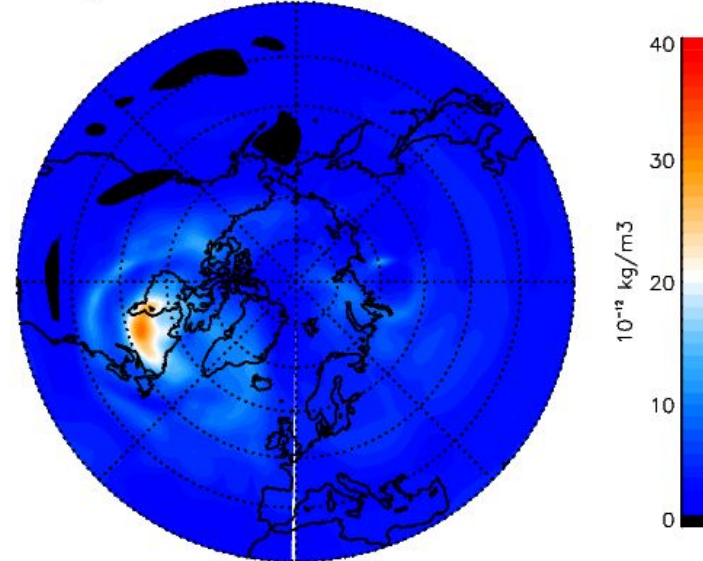


Neutral Density: UT1030

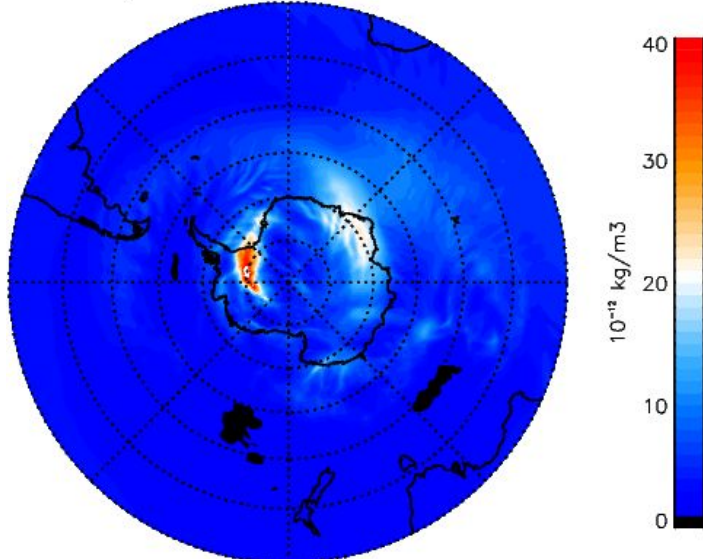
Density 400km 8/24/2005 UT 10:30



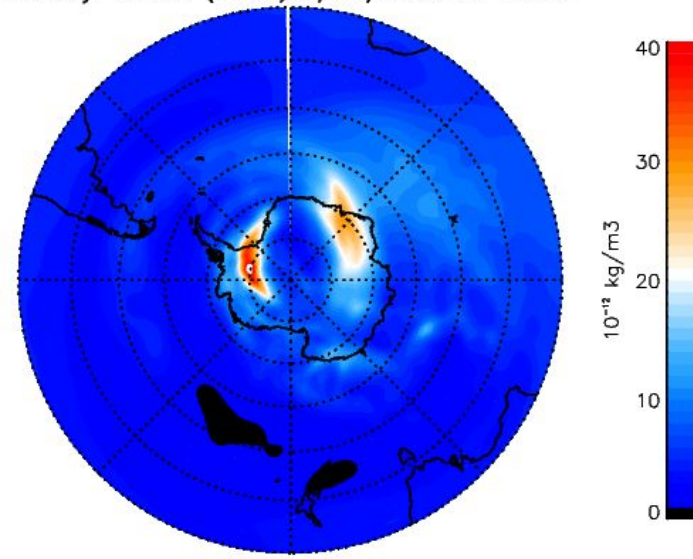
Density 400km (NE30) 8/24/2005 UT 10:30



Density 400km 8/24/2005 UT 10:30

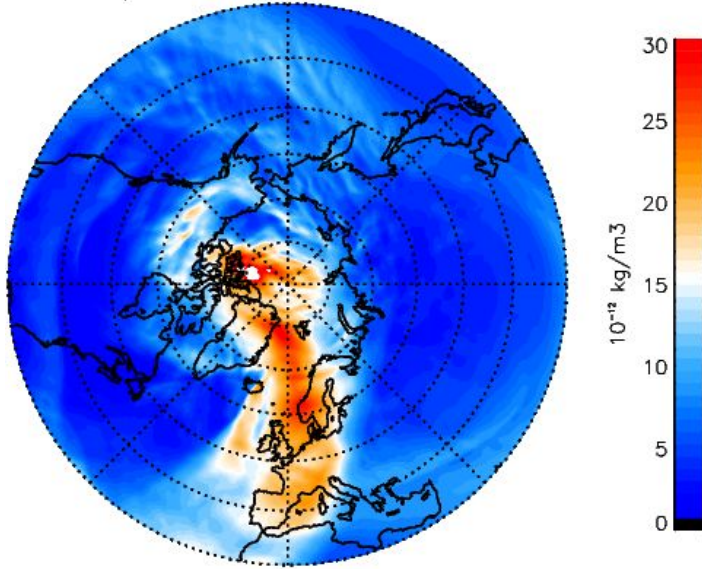


Density 400km (NE30) 8/24/2005 UT 10:30

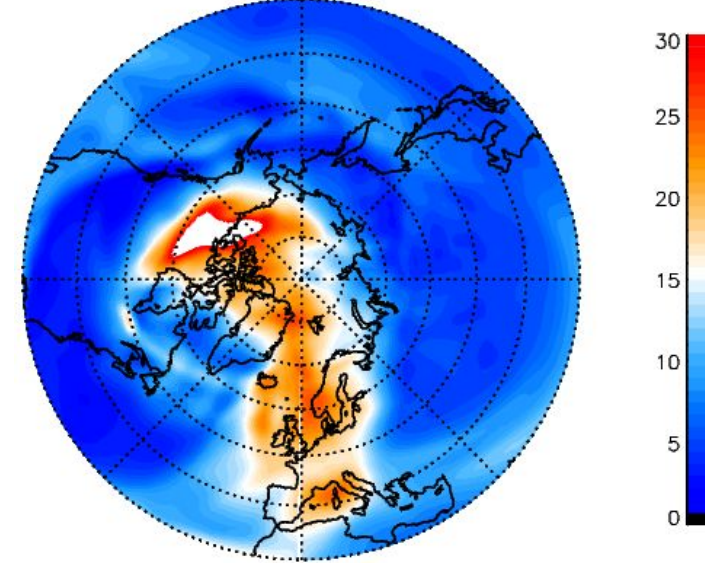


Neutral Density: UT1132

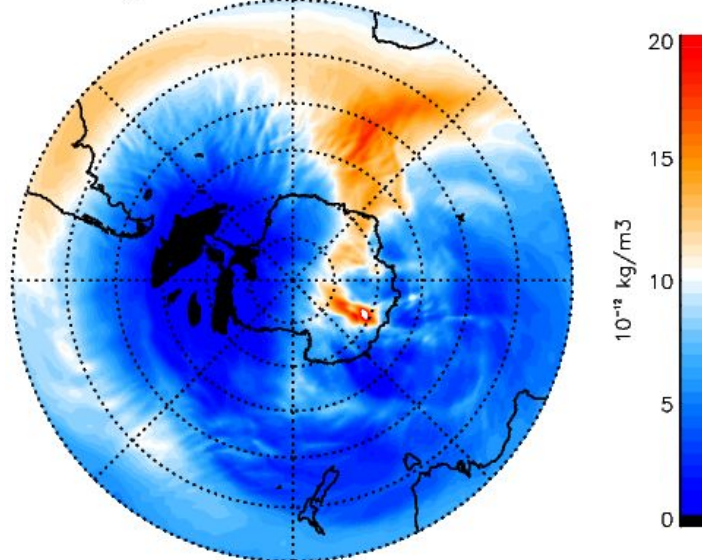
Density 400km 8/24/2005 UT 11:32



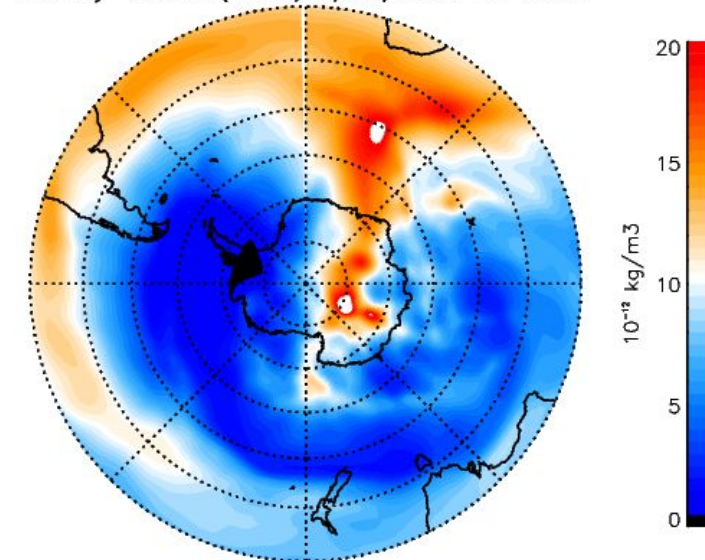
Density 400km (NE30)8/24/2005 UT 11:32



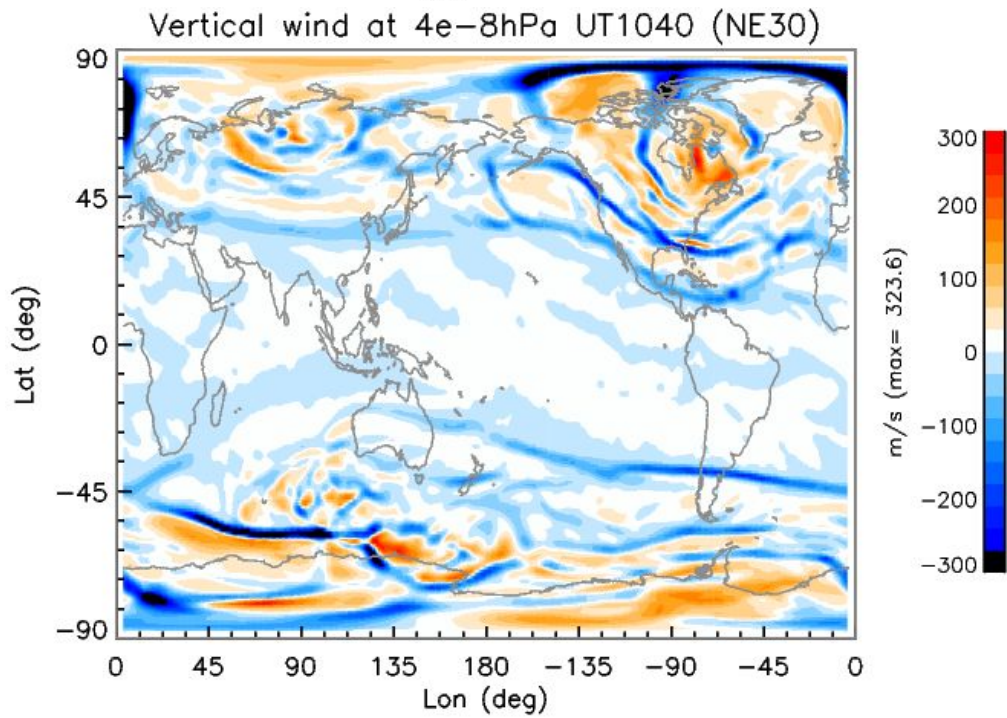
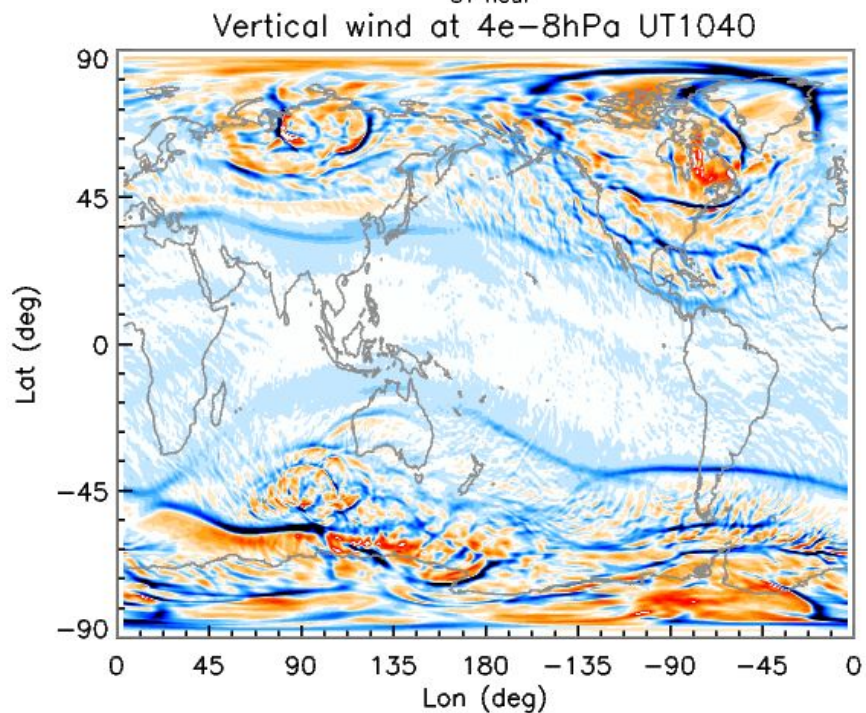
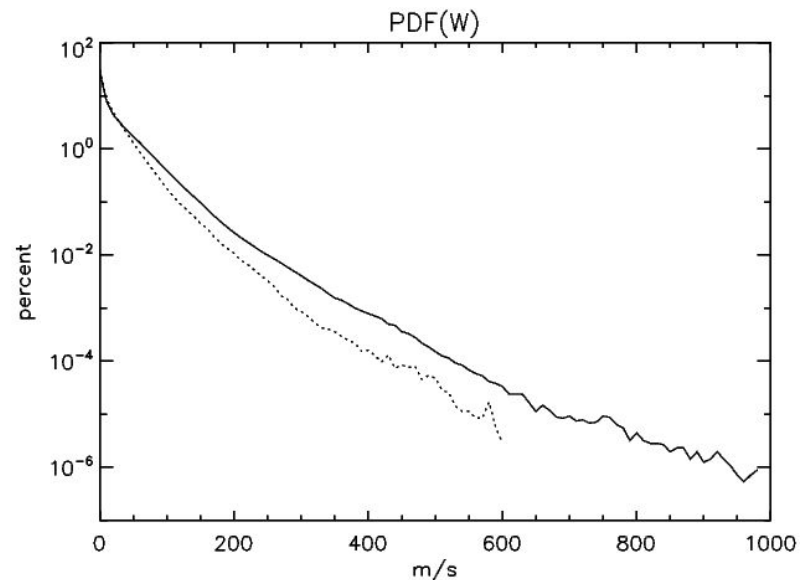
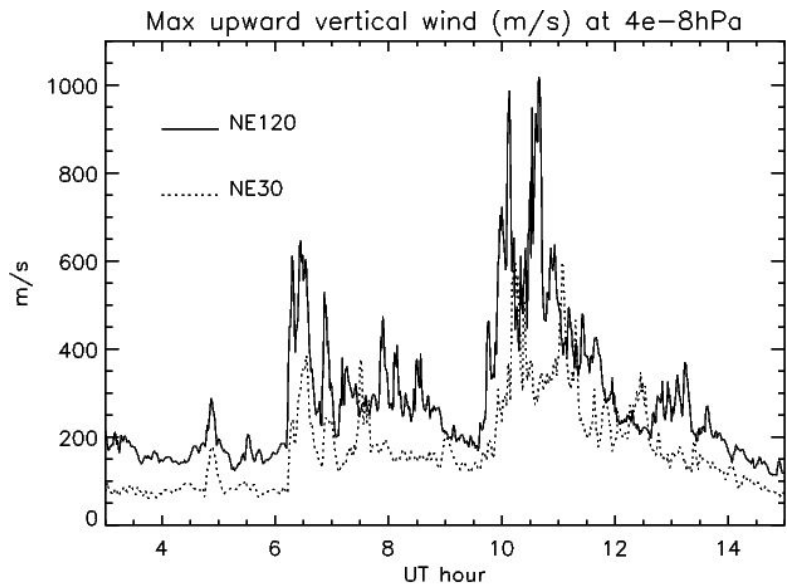
Density 400km 8/24/2005 UT 11:32



Density 400km (NE30) 8/24/2005 UT 11:32



Vertical wind



Summary and Ongoing Work

- In response to the high latitude forcing from GAMERA, the energetics, dynamics, and transport display clear resolution dependence.
- Upper thermosphere Joule heating rates are generally larger in the high-resolution simulation, with larger electric conductivities and electric field perturbations.
- Much stronger tongue of ionization (TOI) is seen in the high-resolution simulation, suggesting stronger plasma transport, which leads to regional enhancement of electric conductivities.
- Stronger dynamical adjustment is seen in high-resolution simulations, both in terms of vertical motion and mesoscale wave structures.
- *Feedback effect to magnetosphere not accounted for.*
- *File-based two way WACCM-X/GAMERA coupling being tested.*